

The Northern California Physics GRE Bootcamp

Held at UC Davis, August 13-14, 2016

Damien Martin

Big picture (your exam)

TOTAL SCORE					
Raw Score	Scaled Score	%	Raw Score	Scaled Score	%
85-100	990	98	43	690	57
84	980	97	41-42	680	54
82-83	970	97	40	670	53
81	960	96	38-39	660	50
80	950	95	37	650	48
78-79	940	95	35-36	640	45
77	930	94	34	630	44
75-76	920	92	33	620	41
74	910	91	31-32	610	39
73	900	90	30	600	37
71-72	890	89	28-29	590	34
70	880	88	27	580	32
68-69	870	87	26	570	29
67	860	86	24-25	560	27
65-66	850	84	23	550	25
64	840	83	21-22	540	22
63	830	82	20	530	20
61-62	820	81	18-19	520	18
60	810	79	17	510	16
58-59	800	78	16	500	13
57	790	76	14-15	490	11
55-56	780	74	13	480	10
54	770	72	11-12	470	7
53	760	71	10	460	6
51-52	750	69	8-9	450	5
50	740	67	7	440	4
48-49	730	65	6	430	3
47	720	63	4-5	420	1
45-46	710	61	3	410	1
44	700	59	1-2	400	1
			0	390	1

*The percent scoring below the scaled score is based on the performance of 10,947 examinees who took the Physics Test between July 1, 2000, and June 30, 2003.

QUESTION Number	Answer	P +	TOTAL C	I
1	C	54		
2	D	30		
3	D	71		
4	C	62		
5	D	28		
6	E	34		
7	B	89		
8	D	65		
9	A	63		
10	A	53		
11	A	28		
12	E	40		
13	B	42		
14	C	27		
15	A	68		
16	D	14		
17	B	81		
18	A	45		
19	B	36		
20	E	49		
21	B	60		
22	A	54		
23	C	45		
24	C	86		
25	E	48		
26	C	30		
27	A	82		
28	E	61		
29	C	63		
30	A	44		
31	A	53		
32	D	62		
33	D	31		
34	C	23		
35	E	82		
36	E	70		
37	D	36		
38	D	35		
39	D	45		
40	D	40		
41	E	66		
42	C	64		
43	D	39		
44	D	54		
45	B	50		
46	E	29		
47	B	46		
48	C	57		
49	E	61		
50	B	50		

QUESTION Number	Answer	P +	TOTAL C	I
51	B	45		
52	C	12		
53	B	32		
54	C	77		
55	E	62		
56	D	54		
57	A	68		
58	B	58		
59	B	87		
60	D	55		
61	C	18		
62	A	35		
63	D	52		
64	A	56		
65	D	44		
66	D	33		
67	E	19		
68	E	51		
69	B	26		
70	B	53		
71	D	32		
72	E	39		
73	D	43		
74	D	50		
75	E	57		
76	C	49		
77	E	44		
78	E	52		
79	D	69		
80	D	28		
81	B	50		
82	D	16		
83	C	30		
84	D	26		
85	A	25		
86	E	24		
87	A	42		
88	C	42		
89	E	37		
90	A	33		
91	B	41		
92	E	45		
93	C	42		
94	E	29		
95	A	42		
96	E	13		
97	E	20		
98	A	72		
99	D	20		
100	B	72		

(24)

20-35 percentile

(6)

<20 percentile

Big picture (1993-1996)

TOTAL SCORE					
Raw Score	Scaled Score	%	Raw Score	Scaled Score	%
67-99	990	97	30-31	690	59
65-66	980	96	29	680	57
64	970	96	28	670	55
63	960	95	27	660	53
62	950	95	26	650	50
61	940	94	24-25	640	48
59-60	930	93	23	630	45
58	920	92	22	620	43
57	910	91	21	610	40
56	900	91	19-20	600	38
54-55	890	90	18	590	35
53	880	89	17	580	32
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46	820	82	10	520	18
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35	730	67			
34	720	65			
33	710	63			
32	700	61			

QUESTION			TOTAL
Number	Answer	P +	C I
1	B	73	
2	B	29	
3	B	55	
4	A	34	
5	B	29	
6	B	43	
7	A	22	
8	A	37	
9	A	40	
10	B	47	
11	D	36	
12	C	36	
13	B	37	
14	D	66	
15	E	12	
16	B	20	
17	E	40	
18	C	77	
19	B	17	
20	D	20	
21	C	27	
22	C	26	
23	D	24	
24	D	70	
25	E	38	
26	C	13	
27	D	49	
28	D	40	
29	E	58	
30	C	28	
31	E	65	
32	B	41	
33	C	56	
34	D	31	
35	A	79	
36	D	46	
37	D	53	
38	D	39	
39	E	54	
40	A	21	
41	A	28	
42	B	52	
43	C	17	
44	A	48	
45	E	45	
46	C	36	
47	C	28	
48	C	30	
49	C	22	
50	C	33	

QUESTION			TOTAL
Number	Answer	P +	C I
51	B	70	
52	C	15	
53	C	34	
54	B	18	
55	A	30	
56	C	42	
57	C	41	
58	E	22	
59	B	36	
60	B	9	
61	B	26	
62	C	13	
63	A	56	
64	C	26	
65	D	44	
66	E	25	
67	C	28	
68	E	61	
69	A	14	
70	D	14	
71	A	20	
72	A	29	
73	C	34	
74	B	21	
75	D	27	
76	B	52	
77	D	12	
78	E	36	
79	D	22	
80	C	31	
81	B	27	
82	E	15	
83	D	15	
84	D	20	
85	B	15	
86	B	36	
87	A	6	
88	B	57	
89	D	18	
90	**	**	
91	E	25	
92	D	15	
93	D	26	
94	D	28	
95	E	23	
96	A	28	
97	E	11	
98	D	39	
99	B	44	
100	C	51	

(35)

20-35 percentile

(18)

<20 percentile

*Percentage scoring below the scaled score is based on the performance of 11,322 examinees who took the Physics Test between October 1, 1993, and September 30, 1996.

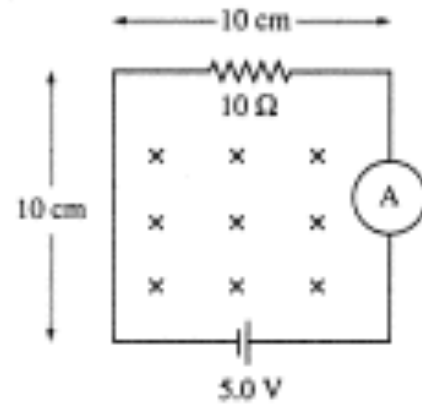
Big tips and tricks

- * Multiple passes through the exam
- * Dimensional analysis (which answers make sense?)
Other hint -- look at exponentials, sines, cosines, ...
- * Expansions, in particular $(1+x)^n = 1 + nx + \dots$
- * Limiting cases (e.g. make parameters go to 0 or infinity)
- * Special cases (e.g. looking at circles)
- * Powers of ten estimation
- * Know scales of things [wavelength / freq of visible light, binding energies of nuclei, mass ratios of common particles (up to muon, pion), ..., mass of stars, mass of galaxies, ...]

Okay to specialize on scales

Big tips and tricks -- material

- * Know your “first year” general physics really well
 - Newtonian mechanics in particular
- * Know your “modern physics” general physics really well (usually Sophomore class)
- * Worth going through
 - Griffiths: Intro to electromagnetism*
 - Griffiths: Intro to quantum mechanics*
 - (Concentrate on harmonic osc, infinite square well, spin systems, expectation values)
 - Schroeder: Thermal physics*
- * Look at the archive of monthly problems in *The Physics Teacher* (if you have access to a university library)



2. The circuit shown above is in a uniform magnetic field that is into the page and is decreasing in magnitude at the rate of 150 tesla/second. The ammeter reads

(A) 0.15 A
 (B) 0.35 A
 (C) 0.50 A
 (D) 0.65 A
 (E) 0.80 A

29

GO ON TO THE NEXT PAGE.

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Evaluate whether
 question is “special”
 or “first year”

67. A black hole is an object whose gravitational field is so strong that even light cannot escape. To what approximate radius would Earth (mass = 5.98×10^{24} kilograms) have to be compressed in order to become a black hole?

(A) 1 nm
 (B) 1 μ m
 (C) 1 cm
 (D) 100 m
 (E) 10 km

28

61

Specific tips and tricks -- material

Bare bones thermal:

Know the definition of partition function, probability of a state, how to find an expectation value

$$Z = \sum_{\text{state } i} \exp(-\beta E_i) = \sum_{\text{energies } j} g_j \exp(-\beta E_j)$$

$$\text{Prob}(\text{State } i) = \frac{1}{Z} e^{-\beta E_i}$$

$$\langle X \rangle = \sum_{\text{state } i} X_i P(X_i) = \frac{1}{Z} \sum_{\text{state } i} X_i \exp(-\beta E_i)$$

Bare bones quantum:

Know the “modern physics” course, know how to find probability of a state, know how to find expectation values, know the special systems, know spin-addition rules

*special systems: particle in box, harmonic oscillator,
two spin 1/2 particles*

Specific tips and tricks -- material

Bare bones (advanced) classical:

Know how to find the Hamiltonian, Hamilton's equations, Lagrangian, and the Euler Lagrange equations for a particle in a gravitational field, charged particle in a uniform electric field, and a pendulum.

Don't just memorize the results, if you can do these three systems you will be aware of the pattern.

Remember: GRE questions are typically short -- cannot get you to do any crazy calculations!

$$C = 3kN_A \left(\frac{h\nu}{kT} \right)^2 \frac{e^{h\nu/kT}}{(e^{h\nu/kT} - 1)^2}$$

65. Einstein's formula for the molar heat capacity C of solids is given above. At high temperatures, C approaches which of the following?

(A) 0

(B) $3kN_A \left(\frac{h\nu}{kT} \right)$

(C) $3kN_A h\nu$

(D) $3kN_A$

(E) $N_A h\nu$

Dimensions here

$$C = 3kN_A \left(\frac{h\nu}{kT} \right)^2 \frac{e^{h\nu/kT}}{(e^{h\nu/kT} - 1)^2}$$

This quantity is dimensionless

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(A) 0

(B) $3kN_A \left(\frac{h\nu}{kT} \right)$

~~(C) $3kN_A h\nu$~~

(D) $3kN_A$

~~(E) $N_A h\nu$~~

Call $x = \frac{h\nu}{kT}$

$$\begin{aligned} C &= 3kN_A x^2 \frac{e^x}{(e^x - 1)^2} \\ &= 3kN_A x^2 \left[\frac{1 + x + \dots}{((1 + x + \dots) - 1)^2} \right] \\ &= 3kN_A x^2 \left[\frac{1}{x^2} + \dots \right] \\ &= 3kN_A + \dots \end{aligned}$$

A distant galaxy is observed to have its H-beta line shifted to a wavelength of 480nm from its laboratory value of 434nm. Which is the best approximation to the velocity of the galaxy?
(Note: $480/434 \sim 1.1$)

- a) $0.01c$
- b) $0.05c$
- c) $0.1c$
- d) $0.32c$
- e) $0.5c$

A distant galaxy is observed to have its H-beta line shifted to a wavelength of 480nm from its laboratory value of 434nm. Which is the best approximation to the velocity of the galaxy?
(Note: $480/434 \sim 1.1$)

- a) $0.01c$
- b) $0.05c$
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$$\lambda_{\text{obs}} = \lambda_{\text{emit}} \sqrt{\frac{c+v}{c-v}}$$

$$\frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}} = \sqrt{\frac{1+(v/c)}{1-(v/c)}} \approx \sqrt{(1+(v/c))^2}$$

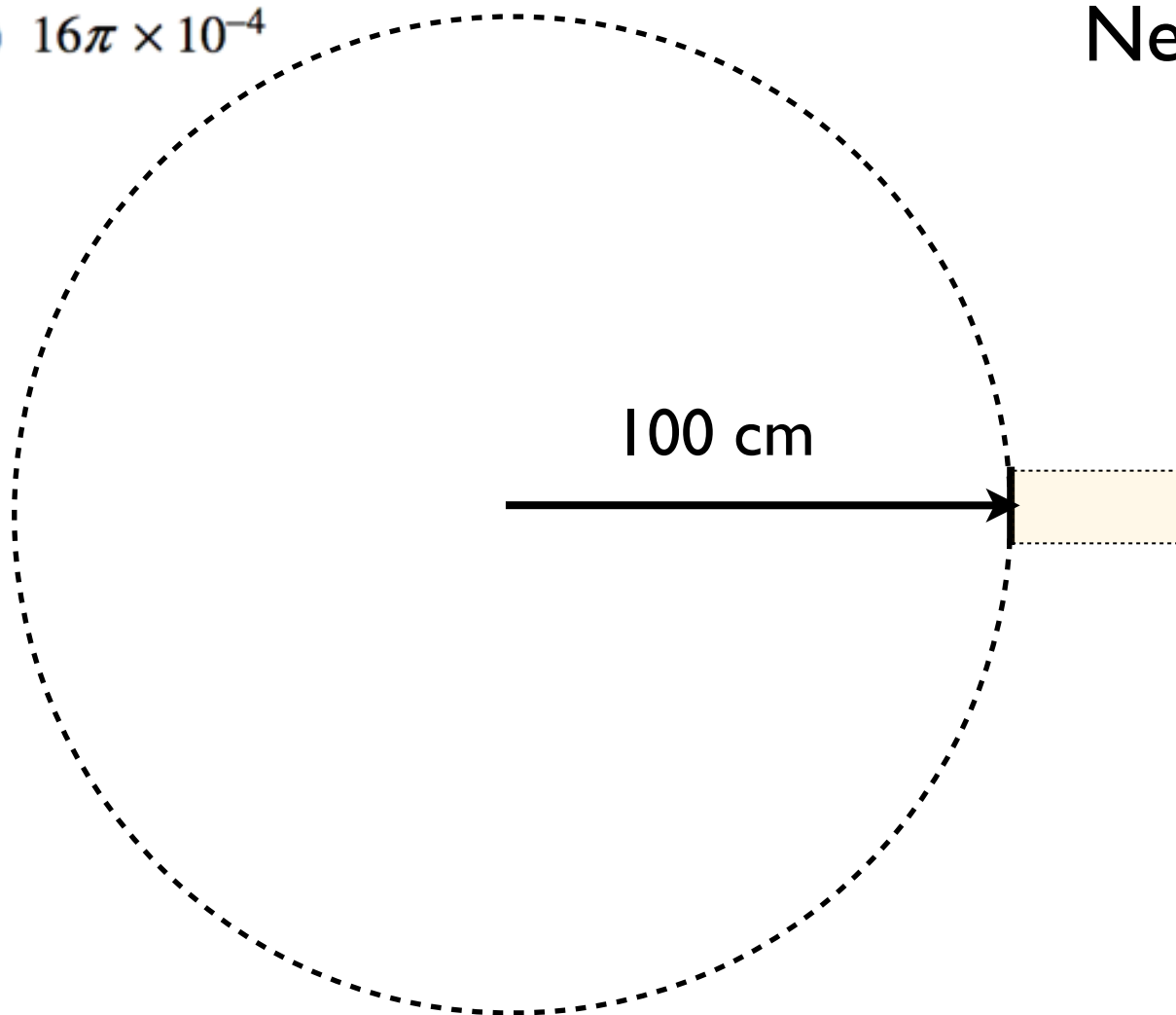
$$v \approx \left(\frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}} - 1 \right) c$$

14. An 8-centimeter-diameter by 8-centimeter-long NaI(Tl) detector detects gamma rays of a specific energy from a point source of radioactivity. When the source is placed just next to the detector at the center of the circular face, 50 percent of all emitted gamma rays at that energy are detected. If the detector is moved to 1 meter away, the fraction of detected gamma rays drops to

- (A) 10^{-4}
(B) 2×10^{-4}
(C) 4×10^{-4}
(D) $8\pi \times 10^{-4}$
(E) $16\pi \times 10^{-4}$

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Need fraction of area covered by sensor

$$A_{\text{sensor}} = \pi(4 \text{ cm})^2 = 16\pi \text{ cm}^2$$

$$A_{\text{sphere}} = 4\pi(100 \text{ cm})^2 = 4\pi \times 10^4 \text{ cm}^2$$

$$\frac{A_{\text{sensor}}}{A_{\text{sphere}}} = \frac{16\pi}{4\pi \times 10^4} = 4 \times 10^{-4}$$

2. The coefficient of static friction between a small coin and the surface of a turntable is 0.30. The turntable rotates at 33.3 revolutions per minute. What is the maximum distance from the center of the turntable at which the coin will not slide?

- (A) 0.024 m
- (B) 0.048 m
- (C) 0.121 m
- (D) 0.242 m
- (E) 0.484 m

2. The coefficient of static friction between a small coin and the surface of a turntable is 0.30. The turntable rotates at 33.3 revolutions per minute. What is the maximum distance from the center of the turntable at which the coin will not slide?

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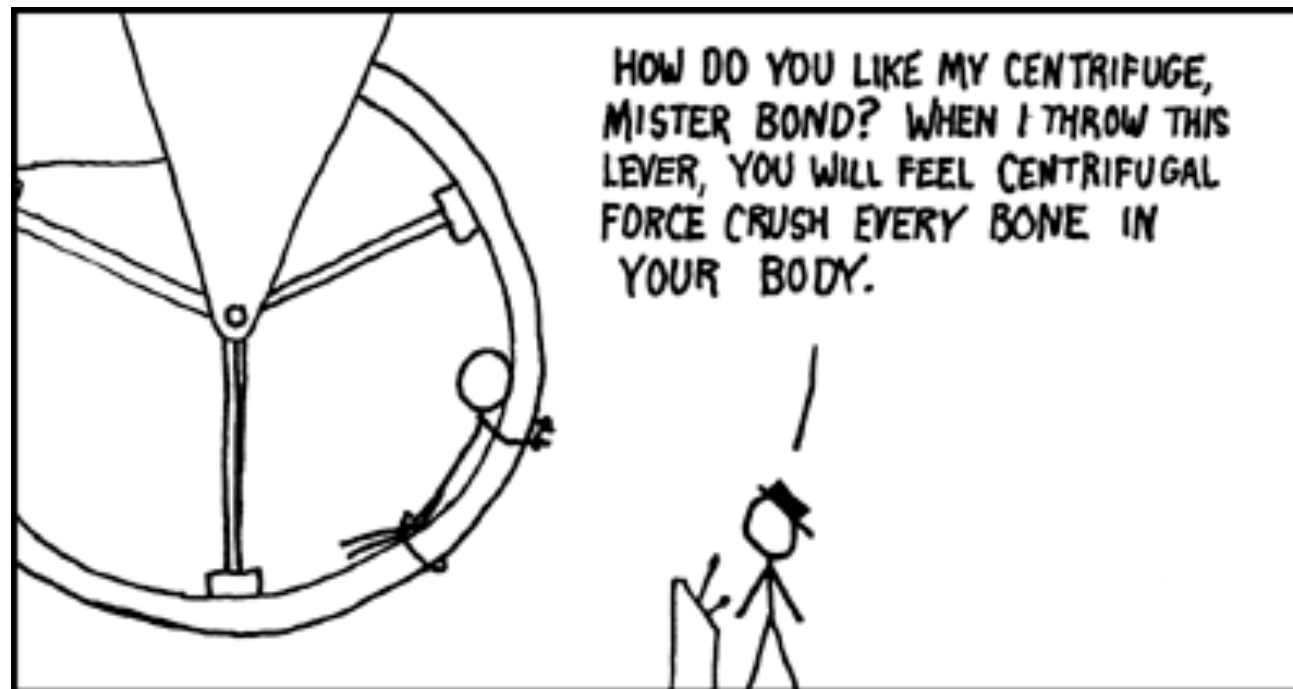
max friction = “centrifugal” force [<http://xkcd.com/123/>]

$$\mu N = \mu mg = m\omega^2 r \quad \text{so} \quad r = \frac{\mu g}{\omega^2}$$

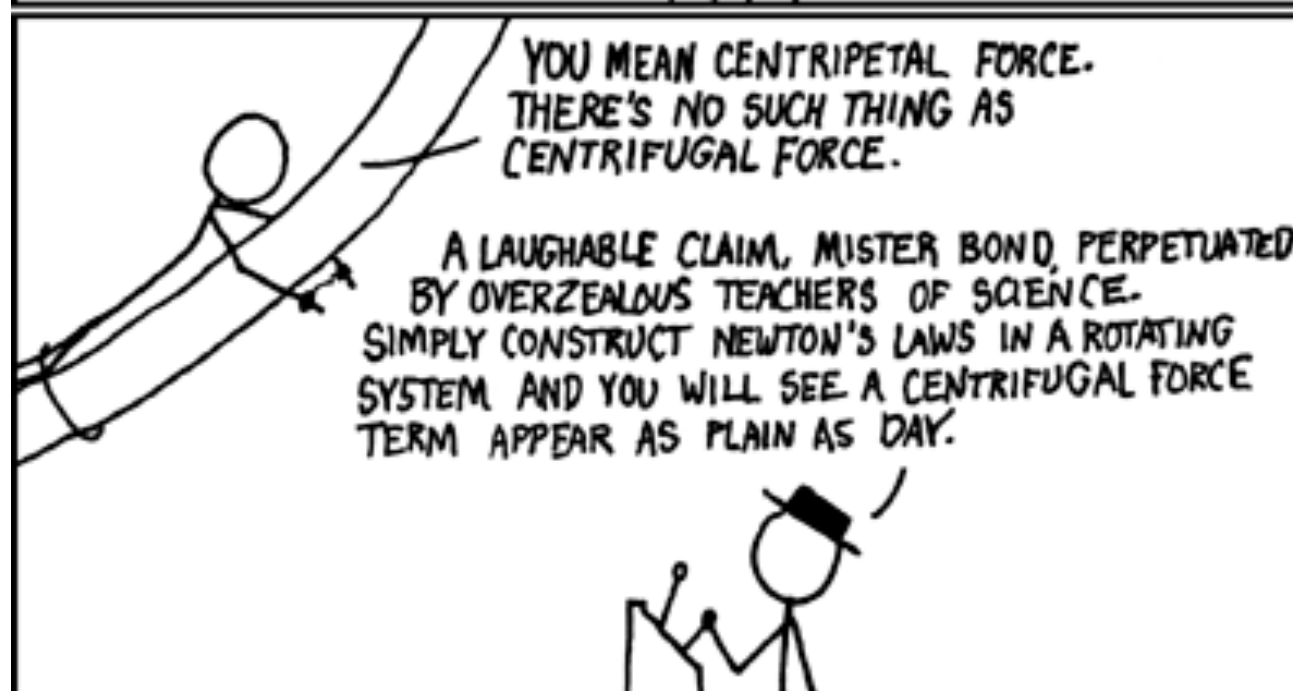
$$\omega = \frac{2\pi}{T} \text{ (don't use!)}$$

$$\omega = \frac{\Delta\theta}{\Delta t} = \frac{2\pi \times 33.3}{60 \text{ s}} \sim \pi \text{ s}^{-1} \quad \Rightarrow \quad \omega^2 \sim 10 \text{ s}^{-2} \text{ therefore}$$

$$r \sim \mu \text{ m} \sim 0.3 \text{ m}$$



HOW DO YOU LIKE MY CENTRIFUGE,
MISTER BOND? WHEN I THROW THIS
LEVER, YOU WILL FEEL CENTRIFUGAL
FORCE CRUSH EVERY BONE IN
YOUR BODY.



YOU MEAN CENTRIPETAL FORCE.
THERE'S NO SUCH THING AS
CENTRIFUGAL FORCE.

A LAUGHABLE CLAIM, MISTER BOND, PERPETUATED
BY OVERZEALOUS TEACHERS OF SCIENCE.
SIMPLY CONSTRUCT NEWTON'S LAWS IN A ROTATING
SYSTEM AND YOU WILL SEE A CENTRIFUGAL FORCE
TERM APPEAR AS PLAIN AS DAY.



COME NOW, DO YOU REALLY EXPECT
ME TO DO COORDINATE SUBSTITUTION
IN MY HEAD WHILE STRAPPED
TO A CENTRIFUGE?

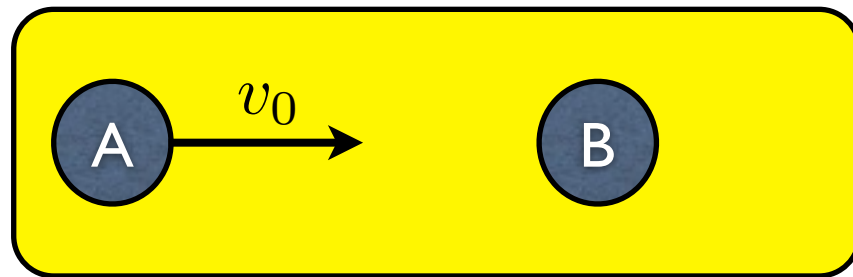
NO, MISTER BOND.
I EXPECT YOU TO DIE.

Things to know (they always seem to come up)

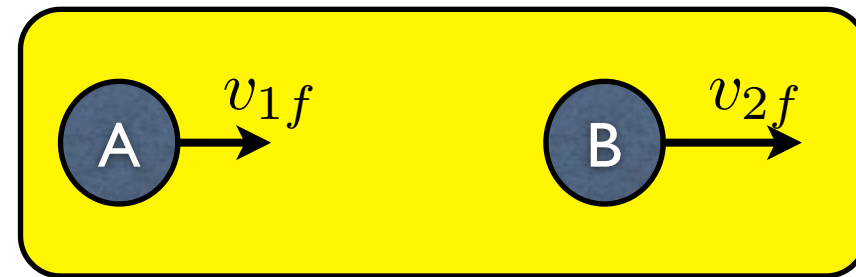
I) Elastic collision formula

Things to know (they always seem to come up)

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Before collision

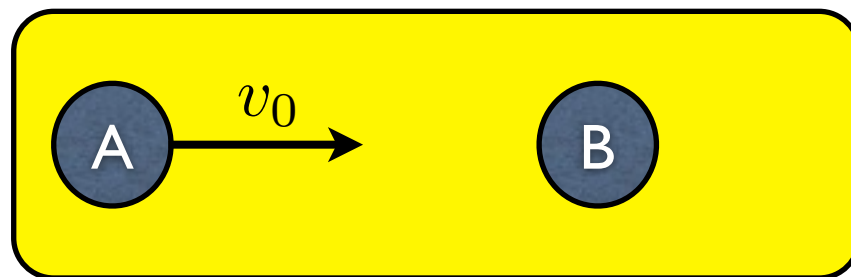


After collision

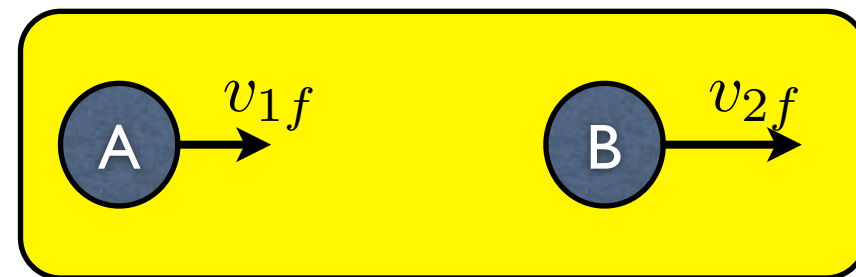
$$v_{1f} = \frac{m_1 - m_2}{m_1 + m_2} v_0$$
$$v_{2f} = \frac{2m_1}{m_1 + m_2} v_0$$

Things to know (they always seem to come up)

1) Elastic collision formula



Before collision



After collision

$$v_{1f} = \frac{m_1 - m_2}{m_1 + m_2} v_0$$
$$v_{2f} = \frac{2m_1}{m_1 + m_2} v_0$$

2) The limiting behavior of capacitors and inductors in DC



acts like



(while uncharged)



(while fully charged)

(e.g. high pass filter question)

3) Virial theorem (and the quick way to get it)

m is reduced mass!

(To get levels for e.g. positronium,
same formula but use reduced mass
for that system)

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$$F(r) = Ar^{+n} \Rightarrow V = \frac{A}{1+n} r^{1+n} = \frac{F(r)}{1+n} r$$

$$\frac{mv^2}{r} = F(r) \Rightarrow \frac{1}{2}mv^2 = \frac{1}{2}F(r)r$$

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4) The Bohr formula (or know how to get it quickly)

$$E = -\frac{Z^2(ke^2)^2m}{2\hbar^2n^2}$$

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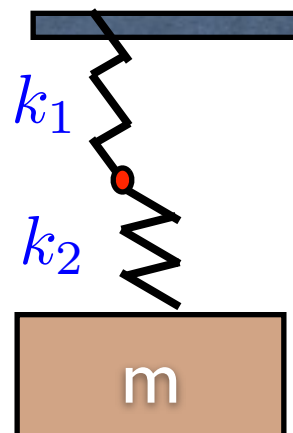
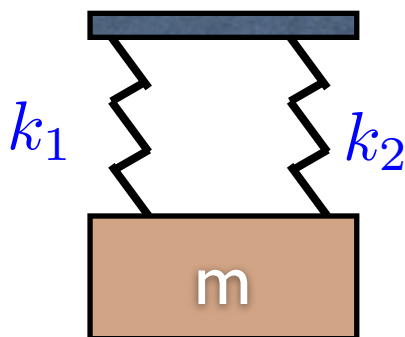
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$$E = -\frac{Z^2 (ke^2)^2 m}{2\hbar^2 n^2}$$

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(To get levels for e.g. positronium, same formula but use reduced mass for that system)

5) Combining masses, springs, capacitors, resistors



Can you find k_{equiv} ?

Frequency of oscillation?

Know reduced mass!

Quick Bohr (semi-classical) derivation

Electron traveling in a circle:

Quick Bohr (semi-classical) derivation

Electron traveling in a circle:

$$\frac{mv^2}{r} = \frac{k(Ze)(e)}{r^2}$$

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Put together to find r (Bohr radius!)

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Put together to find r (Bohr radius!)

$$\frac{1}{r} = \frac{kZe^2}{mv^2r^2} = \frac{kZme^2}{(mvr)^2} = \frac{kZme^2}{n^2\hbar^2}$$

Quick Bohr (semi-classical) derivation

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Potential energy:

Quick Bohr (semi-classical) derivation

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Put together to find r (Bohr radius!)

$$\frac{1}{r} = \frac{kZe^2}{mv^2r^2} = \frac{kZme^2}{(mvr)^2} = \frac{kZme^2}{n^2\hbar^2}$$

Potential energy:

$$PE = -\frac{kZe^2}{r} = -\frac{k^2Z^2me^4}{n^2\hbar^2}$$

Quick Bohr (semi-classical) derivation

Electron traveling in a circle:

$$\frac{mv^2}{r} = \frac{k(Ze)(e)}{r^2}$$

Angular momentum is quantized:

$$L = pr = mvr = n\hbar$$

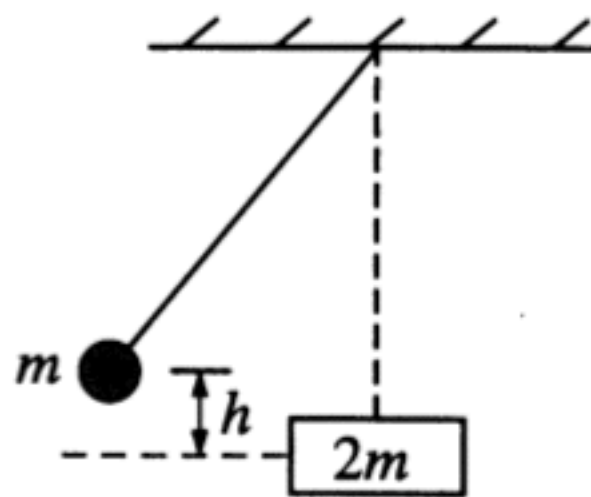
Put together to find r (Bohr radius!)

$$\frac{1}{r} = \frac{kZe^2}{mv^2r^2} = \frac{kZme^2}{(mvr)^2} = \frac{kZme^2}{n^2\hbar^2}$$

Potential energy:

$$PE = -\frac{kZe^2}{r} = -\frac{k^2Z^2me^4}{n^2\hbar^2}$$

Virial thm: $\langle E \rangle = -\langle PE \rangle / 2$



7. As shown above, a ball of mass m , suspended on the end of a wire, is released from height h and collides elastically, when it is at its lowest point, with a block of mass $2m$ at rest on a frictionless surface. After the collision, the ball rises to a final height equal to

- (A) $\frac{1}{9} h$
- (B) $\frac{1}{8} h$
- (C) $\frac{1}{3} h$
- (D) $\frac{1}{2} h$
- (E) $\frac{2}{3} h$

22

Apply elastic collision equations!

66. A sample of radioactive nuclei of a certain element can decay only by γ -emission and β -emission. If the half-life for γ -emission is 24 minutes and that for β -emission is 36 minutes, the half-life for the sample is

- (A) 30 minutes
- (B) 24 minutes
- (C) 20.8 minutes
- (D) 14.4 minutes
- (E) 6 minutes

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$$t_{1/2} \stackrel{?}{=} t_{\gamma} + t_{\beta} \quad \text{or} \quad \frac{1}{t_{1/2}} \stackrel{?}{=} \frac{1}{t_{\gamma}} + \frac{1}{t_{\beta}}$$

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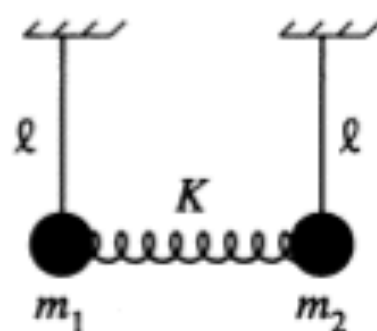
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$$\frac{1}{t_{1/2}} = \frac{1}{t_\gamma} + \frac{1}{t_\beta} = \frac{1}{24} + \frac{1}{36}$$

$$= \frac{1}{6} \left(\frac{1}{4} + \frac{1}{6} \right)$$

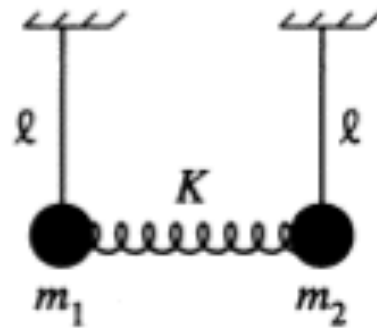
$$= \frac{1}{6} \times \frac{10}{24}$$

$$\Rightarrow t_{1/2} = \frac{24 \times 6}{10} = \frac{144}{10} = 14.4 \text{ min}$$



84. Two pendulums are attached to a massless spring, as shown above. The arms of the pendulums are of identical lengths ℓ , but the pendulum balls have unequal masses m_1 and m_2 . The initial distance between the masses is the equilibrium length of the spring, which has spring constant K . What is the highest normal mode frequency of this system?

- (A) $\sqrt{g/\ell}$
- (B) $\sqrt{\frac{K}{m_1 + m_2}}$
- (C) $\sqrt{\frac{K}{m_1} + \frac{K}{m_2}}$
- (D) $\sqrt{\frac{g}{\ell} + \frac{K}{m_1} + \frac{K}{m_2}}$
- (E) $\sqrt{\frac{2g}{\ell} + \frac{K}{m_1 + m_2}}$



$$f_{\text{high}} \xrightarrow{K \text{ large}} ??$$

$$f_{\text{high}} \xrightarrow{K \text{ small}} ??$$

Reduced mass

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What is the emission energy from a photon going from $n = 3$ to $n = 1$ in *positronium* (one electron and one positron orbiting one another)?

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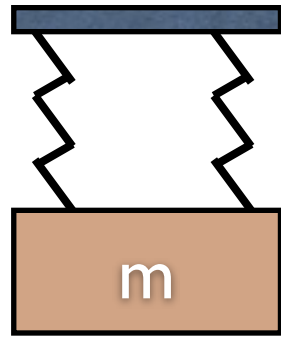
Rule for Hydrogen like atoms: $E_n = \frac{-13.6 \text{ eV}}{n^2} Z^2$

But 13.6 is proportional to the *reduced* mass $m = m_{\text{electron}}$ in Hydrogen

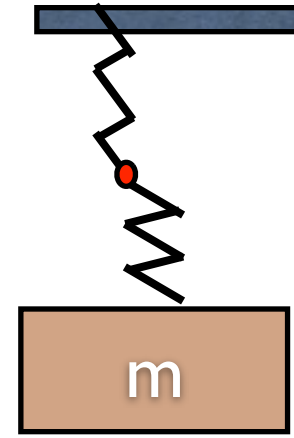
In positronium $m = m_e/2$, so we have to halve the 13.6

$$\boxed{E_n = -\frac{6.8 \text{ eV}}{n^2}}, \quad (\text{positronium energy levels})$$

$$E_{\text{photon}} = E_3 - E_1 = 6.8 \text{ eV} \left(\frac{1}{1^2} - \frac{1}{3^2} \right) = 6.8 \text{ eV} \times \frac{8}{9} \approx 6 \text{ eV}$$



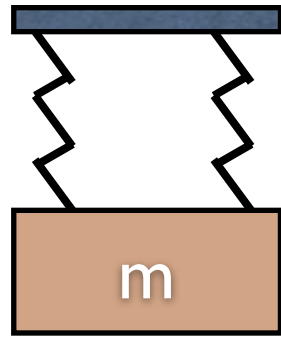
Situation 1)



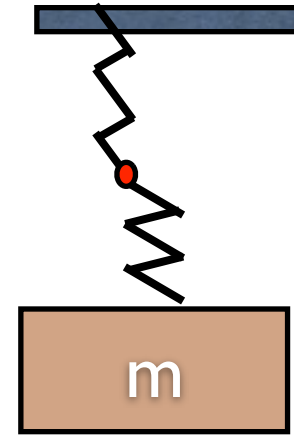
Situation 2)

Two different ways of connecting a mass m to two *identical springs* with spring constant k are shown above. If we denote the frequency of oscillation in situation 1 by f_1 and the frequency of oscillation in situation 2 by f_2 then f_1 / f_2 is:

- a) 4
- b) 2
- c) $1/2$
- d) $1/4$
- e) depends on m and / or k



Situation 1)



Situation 2)

Two different ways of connecting a mass m to two *identical springs* with spring constant k are shown above. If we denote the frequency of oscillation in situation 1 by f_1 and the frequency of oscillation in situation 2 by f_2 then f_1 / f_2 is:

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c) 1/2

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e) depends on m and / or k

(Hint: can pretend k_1 and k_2 are not the same to take limits to determine formula for k_{eff})

$$\frac{f_1}{f_2} = \frac{\omega_1}{\omega_2} = \sqrt{\frac{k_{\text{eff},1}/m}{k_{\text{eff},2}/m}} = \sqrt{\frac{k_{\text{eff},1}}{k_{\text{eff},2}}} = \sqrt{\frac{2k}{k/2}} = 2$$

A particle sits in a periodic potential

$$V(x) = d \sin(kx)$$

What is its oscillation frequency about the minimum?

A particle sits in a periodic potential

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What is its oscillation frequency about the minimum?

Let y be the distance from the minimum. Expanding about the minimum we have:

$$V(y) = V_{\min} + \boxed{0y} + \frac{1}{2} \boxed{\frac{d^2V}{dy^2} \big|_{\min}} y^2 + \dots$$

0 (because min) Just a number, not a function

Force is

$$F = -\frac{dV}{dy} = -\frac{d^2V}{dy^2} \big|_{\min} y + \dots$$

SHM with “spring constant” $k = d^2V/dy^2$ evaluated at min!

spring constant $= -dk^2 \sin(kx) = +dk^2$ evaluated at min

$$f = 2\pi \sqrt{\frac{\text{spring const.}}{m}} = 2\pi \sqrt{\frac{dk^2}{m}}$$

6) Making problems look like a harmonic oscillator

$$\omega^2 = \frac{(d^2V/dx^2)|_{\min}}{m}$$

7) Remember spectroscopic notation (ugh)

$$^{2s+1}(\text{orbital angular momentum symbol})_j$$

and the selection rules for an electric dipole

8) Know the *pattern* of spherical harmonics

Y_ℓ^m

$$Y_0^0(\theta, \varphi) = \frac{1}{2} \sqrt{\frac{1}{\pi}}$$

$(\ell = 0)$

$$Y_1^{-1}(\theta, \varphi) = \frac{1}{2} \sqrt{\frac{3}{2\pi}} \sin \theta e^{-i\varphi}$$

$$Y_1^0(\theta, \varphi) = \frac{1}{2} \sqrt{\frac{3}{\pi}} \cos \theta$$

$$Y_1^1(\theta, \varphi) = \frac{-1}{2} \sqrt{\frac{3}{2\pi}} \sin \theta e^{i\varphi}$$

$(\ell = 1)$

$$Y_2^{-2}(\theta, \varphi) = \frac{1}{4} \sqrt{\frac{15}{2\pi}} \sin^2 \theta e^{-2i\varphi}$$

$$Y_2^{-1}(\theta, \varphi) = \frac{1}{2} \sqrt{\frac{15}{2\pi}} \sin \theta \cos \theta e^{-i\varphi}$$

$$Y_2^0(\theta, \varphi) = \frac{1}{4} \sqrt{\frac{5}{\pi}} (3 \cos^2 \theta - 1)$$

$$Y_2^1(\theta, \varphi) = \frac{-1}{2} \sqrt{\frac{15}{2\pi}} \sin \theta \cos \theta e^{i\varphi}$$

$$Y_2^2(\theta, \varphi) = \frac{1}{4} \sqrt{\frac{15}{2\pi}} \sin^2 \theta e^{2i\varphi}$$

$(\ell = 2)$

Too detailed !

(But if you can remember these, congratulations)

8) Know the *pattern* of spherical harmonics

$$Y_{\ell}^m \quad \begin{array}{l} m - \text{magnetic quantum number } (-\ell, -\ell + 1, \dots, \ell) \\ \ell - \text{orbital quantum number } (0, 1, 2, \dots) \end{array}$$

Y_{ℓ}^m contains φ dependence of the form $e^{im\phi}$

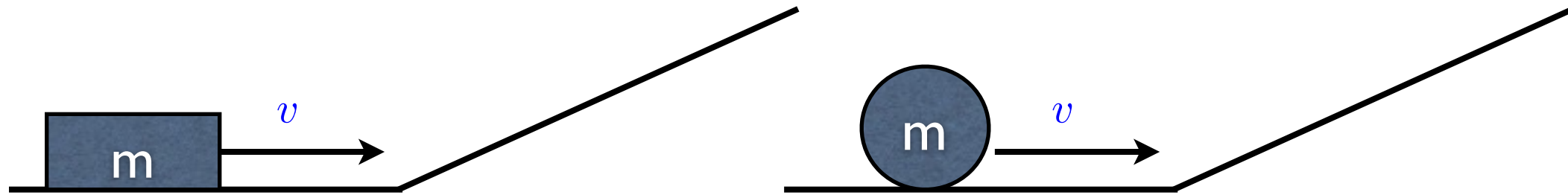
Y_{ℓ}^m contains ℓ dependence of the form $\sin^{\ell} \theta, \sin^{\ell-1} \theta \cos \theta, \dots$

(i.e. can write as ℓ sines or cosines multiplied, or as $\sin(\ell\theta), \cos(\ell\theta)$.)

Compare these rules to the spherical harmonics listed one slide ago.

Random mechanics problem:

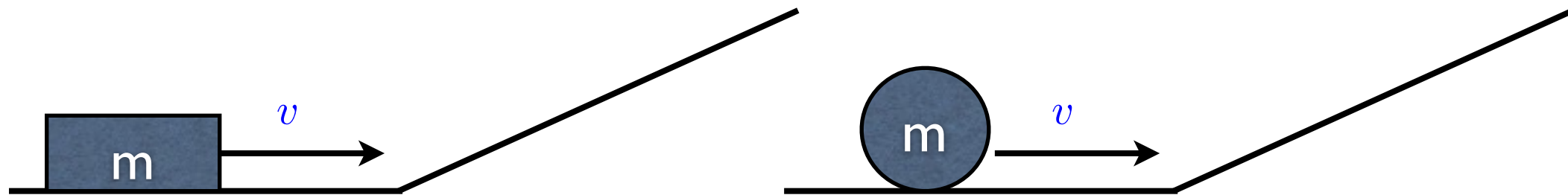
A ball and a block of mass m are moving at the same speed v . When they hit the ramp they both travel up it. The block slides up with (approximately) no friction, the ball experiences just enough friction to roll without slipping. Which goes higher?



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The ball has both translational kinetic energy (equal to that of the block) *and* rotational kinetic energy. Therefore

$$KE_{ball,initial} > KE_{block,initial}$$

The ball converts all this energy into potential energy, and therefore goes higher.